

A VECTOR APPROACH TO THE ALGEBRA OF ROTATIONS WITH APPLICATIONS

GPO PRICE \$ _____

CFSTI PRICE(S) \$ _____

Hard copy (HC) 3.00

Microfiche (MF) .65

653 July 65

BY

PAUL B. DAVENPORT

N 67-18705

FACILITY FORM 602

(ACCESSION NUMBER)
36
(PAGES)
TMX 57187
(NASA CR OR TMX OR AD NUMBER)

(THRU)
1
(CODE)
19
(CATEGORY)

NOVEMBER 1965



GODDARD SPACE FLIGHT CENTER
GREENBELT, MARYLAND



X-546-65-437

A VECTOR APPROACH TO THE ALGEBRA
OF ROTATIONS WITH APPLICATIONS

by

Paul B. Davenport

Goddard Space Flight Center
Greenbelt, Maryland



A VECTOR APPROACH TO THE ALGEBRA
OF ROTATIONS WITH APPLICATIONS

by

Paul B. Davenport
Goddard Space Flight Center

SUMMARY

Under a new type of vector multiplication, a group (whose elements are vectors) is defined which is isomorphic to the group of rotations. This allows a vector representation of rotations which has many advantages over the usual matrix or Eulerian angles approach—e.g., this vector representation avoids the need for trigonometric relationships and requires only three independent parameters.

The simplicity of this vector representation is demonstrated by its use in several applications; in particular, an analytic solution to a least-squares rotation problem is presented. The differential equations defining the motion of a rigid body are also obtained in terms of a vector differential equation.

PRECEDING PAGE BLANK NOT FILMED

CONTENTS

	<u>Page</u>
SUMMARY	iii
1. INTRODUCTION	1
2. VECTORS DEFINING ROTATIONS AND THEIR ALGEBRA	1
3. COORDINATES OF A ROTATED VECTOR	10
4. ROTATIONS DETERMINED BY A VECTOR AND ITS IMAGE	11
5. ROTATIONS DETERMINED BY TWO VECTORS AND THEIR IMAGES	11
6. A LEAST SQUARES ROTATION - DETERMINATION OF ATTITUDE.	15
7. RELATIONS BETWEEN EULERIAN ANGLES, MATRIX OF ROTATION AND VECTOR APPROACH	23
8. EQUATIONS OF MOTION OF A RIGID BODY	26
9. CONCLUSION	31
ACKNOWLEDGMENT	31
REFERENCES	32

PRECEDING PAGE BLANK NOT FILMED.

1. INTRODUCTION

The usual approach to the algebra of rotations is by means of matrix algebra, where the matrix of the rotation is defined by the product of simple rotation matrices. Each of the simple rotations is about a coordinate axis and is uniquely determined by an angle. The most general rotation is uniquely defined by three angles (Eulerian angles) provided that the axis of each rotation is known. Since the convention of defining Eulerian angles varies considerably in the literature, the product matrix is ambiguous when only three angles are known. The angular or matrix approach to rotations is complicated further in that it requires the evaluation of trigonometric functions, which necessitates the use of tables or a computer.

The approach here is to define a group, whose elements are three-dimensional vectors, which is isomorphic to the group of rotation matrices. The algebra of rotations can then be represented by the algebra of vectors (dot and cross product plus a new vector product which will be defined later), and does not require the evaluation of trigonometric functions.

A comment on notation: Matrices and vectors will be denoted by capital English letters and their elements by small English letters with subscripts. All vectors are considered to be column vectors, and a superscript T is used to denote the transpose of a vector or matrix. A superscript of -1 represents the inverse matrix. Primes will be used to denote the transform of a vector. Vector notation, including that for the dot and cross product, is used in this paper merely to simplify the algebraic relationships existing among the components of various vectors, and is not intended to provoke physical interpretation. For example, even though two vectors may represent the same physical quantity, they are not regarded as equal unless they are expressed in the same coordinate system. Similarly, the notation $U = V \times W$ may be used (even though V and W are expressed in different coordinate systems) merely to indicate that the components of U are formed from the components of V and W according to the standard equations for cross products. It will be assumed throughout that all coordinate systems are right-handed orthonormal systems.

2. VECTORS DEFINING ROTATIONS AND THEIR ALGEBRA

If R is the matrix of a rotation then [1] and [2] show that R may be expressed by

$$R_x(\theta) = \cos \theta \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} + (1 - \cos \theta) \begin{pmatrix} x_1^2 & x_1 x_2 & x_1 x_3 \\ x_1 x_2 & x_2^2 & x_2 x_3 \\ x_1 x_3 & x_2 x_3 & x_3^2 \end{pmatrix} + \sin \theta \begin{pmatrix} 0 & x_3 & -x_2 \\ -x_3 & 0 & x_1 \\ x_2 & -x_1 & 0 \end{pmatrix} \quad (1)$$

where $X = (x_1, x_2, x_3)^T$ is a unit vector defining the axis of rotation and θ is the angle of rotation. Unless otherwise stated it is assumed here that X has been selected so that $0 \leq \theta \leq \pi$. This can always be done since $R_x(\theta) = R_{-x}(-\theta)$ and $R_x(\pi + \theta) = R_{-x}(\pi - \theta)$.

The following expression for R may also be found in [1]:

$$R = (B^T)^{-1} B, \quad (2)$$

where

$$B = \begin{pmatrix} 1 & x_3 \tan \frac{\theta}{2} & -x_2 \tan \frac{\theta}{2} \\ -x_3 \tan \frac{\theta}{2} & 1 & x_1 \tan \frac{\theta}{2} \\ x_2 \tan \frac{\theta}{2} & -x_1 \tan \frac{\theta}{2} & 1 \end{pmatrix}. \quad (3)$$

In order to eliminate the angle in equations (1) and (3) we assign a length (which is a function of θ) to X , in such a way that the sine and cosine of θ can be determined by this length. Many such functions exist, and we choose $\tan \theta/2$ and $\sin \theta/2$ as examples. Thus, let

$$Y = \tan \frac{\theta}{2} X, \quad Z = \sin \frac{\theta}{2} X, \quad \text{where } 0 \leq \theta \leq \pi;$$

by standard trigonometric identities we obtain:

$$\sin \frac{\theta}{2} = \frac{\sqrt{Y^2}}{\sqrt{1+Y^2}} = \sqrt{Z^2},$$

$$\cos \frac{\theta}{2} = \frac{1}{\sqrt{1+Y^2}} = \sqrt{1-Z^2},$$

$$\sin \theta = \frac{2\sqrt{Y^2}}{1+Y^2} = 2\sqrt{Z^2(1-Z^2)},$$

$$\cos \theta = \frac{1-Y^2}{1+Y^2} = 1-2Z^2.$$

Hence the definition of R as given by equation (1), using Y and Z respectively, becomes

$$R = \frac{1}{1+Y^2} \left[(1-Y^2) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} + 2 \begin{pmatrix} y_1^2 & y_1 y_2 & y_1 y_3 \\ y_1 y_2 & y_2^2 & y_2 y_3 \\ y_1 y_3 & y_2 y_3 & y_3^2 \end{pmatrix} + 2 \begin{pmatrix} 0 & y_3 & -y_2 \\ -y_3 & 0 & y_1 \\ y_2 & -y_1 & 0 \end{pmatrix} \right], \quad (4)$$

and

$$R = (1-Z^2) \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1 & 0 \\ 0 & 0 & 1 \end{pmatrix} + 2 \begin{pmatrix} z_1^2 & z_1 z_2 & z_1 z_3 \\ z_1 z_2 & z_2^2 & z_2 z_3 \\ z_1 z_3 & z_2 z_3 & z_3^2 \end{pmatrix} + 2\sqrt{1-Z^2} \begin{pmatrix} 0 & z_3 & -z_2 \\ -z_3 & 0 & z_1 \\ z_2 & -z_1 & 0 \end{pmatrix}. \quad (5)$$

As a function of Y, equation (3) becomes

$$B = \begin{pmatrix} 1 & y_3 & -y_2 \\ -y_3 & 1 & y_1 \\ y_2 & -y_1 & 1 \end{pmatrix}. \quad (6)$$

Thus every vector Y defines a rotation matrix by equation (4), and every vector Z with $Z^2 \leq 1$ defines a rotation matrix by equation (5). Hence equations (4) and (5) define mappings which map sets of vectors into the set of rotations. We have yet to verify that the mapping is onto, i.e., that for every rotation matrix there exist vectors Y and Z such that equations (4) and (5) define the matrix of the rotation.

Let R be the matrix (with elements r_{ij}) of a rotation. Separating R into symmetric and skew-symmetric parts, we have

$$R = \frac{1}{2} (R + R^T) + \frac{1}{2} (R - R^T);$$

comparing this with equation (5), we get the following relations:

$$3 - 4Z^2 = \sigma, \quad (7)$$

where σ is the trace of R;

$$2\sqrt{1-Z^2}Z = \frac{1}{2} \begin{pmatrix} r_{23} - r_{32} \\ r_{31} - r_{13} \\ r_{12} - r_{21} \end{pmatrix}, \quad (8)$$

$$2z_2 z_3 = \frac{1}{2} (r_{32} + r_{23}),$$

$$2z_1 z_3 = \frac{1}{2} (r_{13} + r_{31}),$$

$$2z_1 z_2 = \frac{1}{2} (r_{21} + r_{12}),$$

} (9)

$$1 - 2(z_2^2 + z_3^2) = r_{11},$$

$$1 - 2(z_1^2 + z_3^2) = r_{22},$$

$$1 - 2(z_1^2 + z_2^2) = r_{33},$$

These last three equations imply

$$\left. \begin{aligned} z_1^2 &= (1 + r_{11} - r_{22} - r_{33})/4, \\ z_2^2 &= (1 + r_{22} - r_{11} - r_{33})/4, \\ z_3^2 &= (1 + r_{33} - r_{11} - r_{22})/4. \end{aligned} \right\} \quad (10)$$

Thus to express R in the form of equation (5), Z must satisfy nine conditions; however, by using the well-known properties of rotation matrices, it is easy to verify that these nine conditions are consistent.

Equations (7) and (8) give a convenient expression for Z, namely,

$$\begin{aligned} Z^2 &= \frac{3 - \sigma}{4}, \\ Z &= \frac{1}{2\sqrt{1 + \sigma}} \begin{pmatrix} r_{23} & -r_{32} \\ r_{31} & -r_{13} \\ r_{12} & -r_{21} \end{pmatrix}, \end{aligned} \quad (11)$$

provided $Z^2 \neq 1$. If $Z^2 = 1$, then equations (9) and (10) may be used to determine Z. In this case, however, Z is not unique, since $-Z$ also satisfies equations (9) and (10) if Z does. In practice, this ambiguity is of no consequence, since if $Z^2 = 1$, Z and $-Z$ yield the same matrix.

From equation (1), we find that $\sigma = 1 + 2 \cos \theta$, which by equation (11) implies that $Z^2 \leq 1$.

Hence the mapping defined by equation (5) is a mapping of three-dimensional vectors over the field of real numbers whose length is less than or equal to unity (denote this set of vectors by ζ) onto the group of rotation matrices. The mapping is also one-to-one except when $Z^2 = 1$ ($\sigma = -1$).

The vector Y may be obtained either directly from the matrix R (in a manner analogous to the preceding), or from the vector Z by using the relationship

$$Y = \frac{1}{\sqrt{1-Z^2}} Z . \quad (12)$$

Thus,

$$Y = \frac{1}{1+\sigma} \begin{pmatrix} r_{23} & -r_{32} \\ r_{31} & -r_{13} \\ r_{12} & -r_{21} \end{pmatrix} . \quad (13)$$

The vector Y is undefined when $\sigma = -1$. This singularity may be removed if we agree to allow vectors of infinite magnitude whose direction is given by a unit vector, X say. For such vectors, equation (4) becomes

$$R = -I + 2XX^T . \quad (14)$$

With this convention, Y may be obtained from the rotation matrix R with trace -1 merely by observing that $X = Z$ in this special case. Hence, X may be obtained from equations (9) and (10).

Let η denote the set of all real three-dimensional vectors augmented by the vectors of infinite magnitude discussed above. Then equations (4) and (14) define a mapping from η onto the group of rotation matrices.

Thus, either of the two sets ζ or η may be used to parametrize the group of rotations. However, when vectors are used for this purpose it should be emphasized that they correspond to transformations. Hence, it is their algebraic properties as transformations that we are primarily concerned with, and not their properties as vectors (indeed ζ is not even a vector space). It is of no consequence that two elements of ζ or η may be equal in the sense of equality of transformations but not equal in the normal sense of vector equality. On the other hand, the abundance of vector operations and their algebraic properties that are normally employed to simplify relationships between components of vectors can also be applied to vectors corresponding to a rotation. Indeed, it is the fact that the algebra of vectors as rotations can be expressed in terms of standard vector notation that makes the vector parametrization appealing. In

order to distinguish vectors used to parametrize rotations from ordinary vectors which are transformed by rotations, we introduce a new type of vector which is obtained by defining a new equivalence relation among real three-dimensional vectors. Definition: Given a set δ of real three-dimensional vectors and a mapping τ which maps δ onto the group of rotation matrices then two elements of δ are said to be equivalent if they map into the same rotation matrix.

For the two sets ζ and η defined above, vector equivalence is the same as vector equality except for vectors mapping into rotation matrices with trace of -1.

Clearly, vector equivalence as defined above is an equivalence relation and separates the sets ζ and η into disjoint classes.

These equivalence classes are merely the inverse images under τ of the rotation matrices. We shall call these equivalence classes "rotation vectors" and denote them in the same manner as ordinary vectors.

The usual algebra associated with vectors is also applicable to rotation vectors. The well-known operations of scalar multiplication, vector addition, dot and cross products are performed on rotation vectors in the classical manner and the ordinary symbolism is used to denote these operations. When there is more than one vector in the equivalence class, the result of an operation is that obtained by using either vector. However, the same vector must be used throughout any one expression.

Given two rotation vectors Y_1 and Y_2 , equation (4) defines two rotation matrices, say R_1 and R_2 respectively. It is well-known that $R = R_2 R_1$ is also the matrix of a rotation. Thus by a previous discussion there exists a rotation vector Y which defines R . This rotation vector Y can be obtained from equation (13) by forming the product $R_2 R_1$ in terms of Y_1 and Y_2 , which gives the remarkably simple expression

$$Y = \frac{1}{1 - Y_1 \cdot Y_2} (Y_1 + Y_2 + Y_1 \times Y_2). \quad (15)$$

If Z_1 and Z_2 are members of ζ defining R_1 and R_2 respectively then the $Z \in \zeta$ defining R , such that $R = R_2 R_1$, may be obtained in a similar manner i.e. by comparing the trace and skew-symmetric part of $R_2 R_1$ with the trace (σ) and skew-symmetric part of R respectively as given by equation (5). This gives

$$\sigma = 3 - 4Z^2 = 3 - 4Z_0^2,$$

$$\sqrt{1-Z^2} Z = \left(\sqrt{1-Z_1^2} \sqrt{1-Z_2^2} - Z_1 \cdot Z_2 \right) Z_0 ,$$

where

$$Z_0 = \sqrt{1-Z_2^2} Z_1 + \sqrt{1-Z_1^2} Z_2 + Z_1 \times Z_2 .$$

Thus

$$1 - Z^2 = 1 - Z_0^2 = \left[\sqrt{1-Z_1^2} \sqrt{1-Z_2^2} - Z_1 \cdot Z_2 \right]^2 ,$$

and

$$\frac{\sqrt{1-Z_1^2} \sqrt{1-Z_2^2} - Z_1 \cdot Z_2}{\sqrt{1-Z^2}} = \pm 1 ,$$

where the sign depends only on the numerator since $\sqrt{1-Z^2} \geq 0$ by definition i.e.

$\sqrt{1-Z^2} = \cos \theta/2$ where $0 \leq \theta \leq \pi$. Hence, using the signum function, $\text{sgn}(x)$, which is defined to be +1, -1 according as x is positive or negative, respectively (here zero is considered positive or negative)

$$Z = \left[\text{sgn} \left(\sqrt{1-Z_1^2} \sqrt{1-Z_2^2} - Z_1 \cdot Z_2 \right) \right] Z_0 \quad (16)$$

The above argument does not guarantee that the Z as given by equation (16) corresponds to $R = R_2 R_1$ when $Z^2 = 1$. However, it is straightforward, but tedious, to verify that equation (16) indeed gives the Z defining $R = R_2 R_1$ for all Z_1 and Z_2 contained in ζ corresponding to R_1 and R_2 respectively.

Actually equation (16) may be obtained more readily from equations (12) and (15), and the identity

$$Z = \frac{1}{\sqrt{1+Y^2}} Y \quad (17)$$

which is a useful formula since it is also valid in the limit as Y^2 approaches infinity.

Equations (15) and (16) suggest a new type of vector product. Definition: Let δ be a set of vectors, let $*$ be a binary operation on δ and let τ be a mapping of δ onto the group of rotation matrices. Then $*$ is said to be a rotation product if it is preserved by τ , i.e. if $\tau(V*W) = \tau(V)\tau(W)$, for every V and W in δ .

Thus for Z_1, Z_2 in ζ , the product $Z = Z_2 * Z_1$ given by equation (16) is a rotation product. In a similar vein, for Y_1, Y_2 in η , equation (15) defines a rotation product, except when $Y_1 \cdot Y_2 = 1$, or when either of the rotation vectors has infinite magnitude. In these exceptional cases, we define the product $Y_2 * Y_1$ by forming Z_1 and Z_2 by equation (17), $Z_2 * Z_1$ by equation (16) and then $Y_2 * Y_1$ by equation (12).

Let η' and ζ' denote the two sets of rotation vectors (equivalence classes) defined by vector equivalence and the sets η and ζ respectively. Then each of the two sets η' and ζ' , together with its rotation product, forms a group.

Clearly η' is closed under the rotation product. Also, since equation (17) implies that $Z = Z_2 * Z_1$ has length less or equal to unity, ζ' is closed. The associativity of the rotation product is easily verified. The identity of each set is the null rotation vector, and the inverse of V (in either set) is $-V$. Furthermore, the respective mappings of the two sets defined herein are isomorphisms onto the group of rotation matrices. Thus I (the identity rotation matrix) is the image of the null rotation vector and R^{-1} is the image of $-V$ if R is the image of V .

Actually, a one-to-one relationship between vectors and rotations may be defined without the concept of equivalence class. For example, when $\sigma = -1$ additional conditions on Z may be imposed so as to insure uniqueness. The choice of which additional constraints to select, however, depends on the preference of the user (a situation analogous to the many definitions of Eulerian angles). Thus, we introduced vector equivalence to emphasize that for the purpose of representing a rotation it does not matter which vector one selects out of an equivalence class for a 180 degree rotation.

The group of rotation transformations can now be represented by either of the two groups ζ' or η' . This representation has certain advantages over the usual matrix representation, stemming from the fact that the rotation is defined by three independent parameters, without recourse to trigonometry (the matrix approach requires nine elements, or the evaluation of six trigonometric functions plus two matrix multiplications). In many applications the vector representation requires fewer calculations than the matrix representation. For example, matrix

multiplication requires 27 scalar multiplications plus 18 additions; the rotation product of equation (15) requires only 13 multiplications and 10 additions.

The choice of which vector representation to use depends upon the application. The formulas associated with the Y rotation vector are in general quite simple, requiring no radicals or trigonometric functions. Thus this vector representation is a valuable tool for hand calculations or for deriving theoretical results, but has the inconvenience of becoming infinite for all 180-degree rotations. The Z vector representation, on the other hand, is always finite, always defined (uniquely, except for its sign at 180-degree rotations), and the rotation product is valid for all rotations. However, the existence of the radical is inconvenient for hand calculations. The vector representation $W = \tan \theta/4X$ may prove to be useful since it combines some of the assets of both the Y and Z. In this case we have

$$R = \frac{1}{(1+W^2)^2} \left[(W^4 - 6W^2 + 1) I + 8WW^T + 4(1-W^2) \begin{pmatrix} 0 & w_3 & -w_2 \\ -w_3 & 0 & w_1 \\ w_2 & -w_1 & 0 \end{pmatrix} \right],$$

$$Y = \frac{2}{1-W^2} W, \quad Z = \frac{2}{1+W^2} W,$$

$$W = \frac{1}{1+\sqrt{1+Y^2}} Y = \frac{1}{1+\sqrt{1-Z^2}} Z.$$

3. COORDINATES OF A ROTATED VECTOR

One of the most frequent uses of a rotation matrix is to find the coordinates of a vector after the rotation. The vector representation yields a useful formula for this application. Let V' be the image of V under the rotation and R the matrix of the rotation (i.e., $V' = RV$). If Y and Z are the rotation vectors defining R , then equations (4), (5), and (14) give V' as

$$V' = YV = \begin{cases} \frac{1}{1+Y^2} \left[(1-Y^2)V + 2(V \cdot Y)Y + 2V \times Y \right], & Y^2 < \infty \\ -V + 2(V \cdot X)X, & Y^2 = \infty \end{cases}$$

$$V' = ZV = (1-Z^2)V + 2(V \cdot Z)Z + 2\sqrt{1-Z^2}V \times Z$$

Note that we have used the symbolism YV to denote the image of V under the rotation corresponding to Y . Thus the above equations define a vector multiplication in the same sense that RV is used to denote a matrix multiplication.

4. ROTATIONS DETERMINED BY A VECTOR AND ITS IMAGE

In many instances, one is given two vectors (V and V') and desires to determine the matrix R such that $V' = RV$. If R is to be a rotation matrix, then we must have $V'^2 = V^2$. For simplicity, we assume that $V'^2 = V^2 = 1$. The rotation taking V into V' is not unique; however, the vector representation of the "shortest path" rotation is immediately obvious—the axis of rotation is collinear with $V' \times V$ and the angle of rotation is $\cos^{-1}(V \cdot V')$. Thus

$$Y = \frac{1}{1 + V \cdot V'} V' \times V, \quad V \cdot V' \neq -1$$

$$Z = \frac{1}{\sqrt{2(1 + V \cdot V')}} V' \times V, \quad V \cdot V' \neq -1.$$

If $V \cdot V' = -1$ then Z is any vector satisfying the two conditions $Z^2 = 1$ and $Z \cdot V = 0$; Y has infinite magnitude with direction defined by Z .

5. ROTATIONS DETERMINED BY TWO VECTORS AND THEIR IMAGES

A common practice in orbit theory is to construct a rotation that takes the xy -plane into the plane of the orbit and also takes the x -axis into the direction of perigee. This is a particular example of the following problem: Given V_1, V_2 (V_1 and V_2 noncollinear), $V_1',$ and V_2' such that $V_1'^2 = V_1^2, V_2'^2 = V_2^2,$ and $V_1' \cdot V_2' = V_1 \cdot V_2,$ find the rotation that takes V_1 into V_1' and V_2 into V_2' . Here again, the vector representation of the rotation yields a simple solution.

If $V' = RV$, then equation (2) implies that $B^T V' = BV$. By equation (6) and matrix multiplication, this condition can be written as

$$V' + Y \times V' = V - Y \times V,$$

or

$$V - V' = Y \times (V + V'),$$

where Y is the vector representation of R .

Thus, the conditions $V_1' = RV_1$ and $V_2' = RV_2$ may be expressed by

$$V_1 - V_1' = Y \times (V_1 + V_1') \quad \text{and} \quad V_2 - V_2' = Y \times (V_2 + V_2').$$

Let

$$A_i = V_i - V_i', \quad B_i = V_i + V_i' \quad (i = 1, 2)$$

then the condition equations become

$$A_i = Y \times B_i, \quad (i = 1, 2)$$

where $A_i \cdot B_i = 0$ and $A_1 \cdot B_2 = -A_2 \cdot B_1$. It is immediately obvious that $Y \cdot A_i = 0$. If each side of the first equation is cross multiplied by the vector A_2 , we obtain

$$A_2 \times A_1 = A_2 \times (Y \times B_1) = (A_2 \cdot B_1) Y - (A_2 \cdot Y) B_1 = (A_2 \cdot B_1) Y;$$

thus if $A_2 \cdot B_1 = -A_1 \cdot B_2 \neq 0$, we obtain a simple expression for Y , namely,

$$Y = \frac{1}{V_1 \cdot V_2' - V_2 \cdot V_1'} (V_1 - V_1') \times (V_2 - V_2'). \quad (18)$$

A more general expression may be obtained as follows. Cross multiply each side of the i^{th} equation by B_i ; this gives

$$B_i \times A_i = B_i^2 Y - (B_1 \cdot Y) B_i \quad (i = 1, 2) \quad (19)$$

Therefore,

$$B_2 \cdot (B_1 \times A_1) = B_1^2 B_2 \cdot Y - (B_1 \cdot B_2) B_1 \cdot Y,$$

$$B_1 \cdot (B_2 \times A_2) = B_2^2 B_1 \cdot Y - (B_1 \cdot B_2) B_2 \cdot Y,$$

and solving these two linear equations for $B_1 \cdot Y$ and $B_2 \cdot Y$ yields

$$B_1 \cdot Y = \frac{(B_1 \times B_2) \cdot [(B_1^2 A_2) - B_1 \cdot B_2 A_1]}{(B_1 \times B_2)^2},$$

$$B_2 \cdot Y = \frac{(B_1 \times B_2) \cdot [(B_1 \cdot B_2) A_2 - B_2^2 A_1]}{(B_1 \times B_2)^2}$$

Substitution of these last expressions into equation (19) gives the solution for Y , provided $B_1 \times B_2 \neq 0$.

If $V_1' = V_1$ and/or $V_2' = V_2$, the Z rotation vector is obtained from Y by equations (17) and (19) unless $B_1 \times B_2 = 0$; otherwise, equations (17) and (18) provide a much simpler expression for Z , namely,

$$Z = \frac{\text{sgn}(V_1 \cdot V_2' - V_2 \cdot V_1')}{\sqrt{(V_1 \cdot V_2' - V_2 \cdot V_1')^2 + [(V_1 - V_1') \times (V_2 - V_2')]^2}} (V_1 - V_1') \times (V_2 - V_2').$$

When $V_i' = V_i$ and $B_1 \times B_2 = 0$, $Z = V_i$.

The above expressions provide another means of determining the vector representation of a rotation from the matrix of the rotation R . Merely choose two independent vectors (V_1 and V_2) and set $V_1' = RV_1$, $V_2' = RV_2$.

Given V_1, V_2, V_1' and V_2' such that $V_1^2 = V_1'^2$, $V_2^2 = V_2'^2$ and $V_1 \cdot V_2 = V_1' \cdot V_2'$, the matrix of the rotation, R , which takes V_1 into V_1' and V_2 into V_2' can be determined from the vector representation obtained above and equation (4) or (5). A more direct approach, however, is to construct a right-handed orthonormal coordinate system from the vectors V_1 and V_2 and a second system from the vectors V_1' and V_2' in the same manner. The rotation taking the first system into the second will then take V_i into V_i' . The matrix of this rotation can be easily written as the product of two matrices, each of whose rows or columns are formed from the components of the constructed axes relative to some underlying coordinate system. Let

$$U_1 = V_1/|V_1|, \quad U_1' = V_1'/|V_1'|,$$

$$W = \left[V_2 - \frac{(V_1 \cdot V_2)}{V_1^2} V_1 \right], \quad W' = \left[V_2' - \frac{(V_1' \cdot V_2')}{V_1'^2} V_1' \right],$$

$$U_2 = W/|W|, \quad U_2' = W'/|W'|,$$

$$U_3 = U_1 \times U_2, \quad U_3' = U_1' \times U_2',$$

$$R_1 = (U_1, U_2, U_3), \quad R_2 = (U_1', U_2', U_3');$$

then

$$R = R_2 R_1^{-1}.$$

6. A LEAST SQUARES ROTATION - DETERMINATION OF ATTITUDE¹

Consider a vehicle with a local coordinate system which has been rotated from a fixed coordinate system. If the vehicle has equipment capable of determining the direction (relative to the local coordinate system) of a point whose direction relative to the fixed coordinate system is known, then the rotation relating the local and fixed systems can be determined by two or more observations of such points.² If the directions were exact, then the rotation relating the two coordinate systems could be obtained from any two noncollinear observations by the methods of the previous section. In practice, however, exactness is not obtainable and in such cases one generally seeks the "least squares" solution.

For the case at hand, we seek a rotation matrix R such that the scalar function

$$\phi(R) = \sum_{i=1}^n (W_i - RV_i)^2$$

is a minimum. Here V_i and W_i are vectors defining the directions of a point relative to the fixed and local coordinate systems, respectively. The matrix R —thus also the function $\phi(R)$ —has only three independent parameters. Any rotation which makes $\phi(R)$ a minimum is either a solution of the three equations obtained by setting the partial derivatives of ϕ with respect to each independent parameter to zero or at a point where the partials do not exist. The use of Eulerian angles as the independent parameters leads to three condition equations which are quite complicated, tedious to derive, and must be dealt with as three scalar equations. The condition equations in terms of a vector representation of R , on the other hand, can be expressed as a single vector equation which is easily derived and has a solution obtainable by vector and matrix algebra.

Since R is a rotation matrix, $\phi(R)$ may be written as

$$\phi(R) = \sum_{i=1}^n \left[V_i^2 + W_i^2 - 2W_i \cdot (RV_i) \right],$$

¹This section is a solution to Problem 65-1 by Grace Wahba in SIAM Review Vol. 7, No. 3, July 1965, which appeared during the writing of this paper.

²The Orbiting Astronomical Observatory has this capability, where the points are known stars.

which as a function of $Y \in \eta$ gives

$$\phi(\mathbf{R}) = \sum_{i=1}^n \left\{ V_i^2 + W_i^2 - \frac{2}{1+Y^2} \left[(1-Y^2) \mathbf{V}_i \cdot \mathbf{W}_i + 2(\mathbf{V}_i \cdot \mathbf{Y})(\mathbf{W}_i \cdot \mathbf{Y}) + 2(\mathbf{W}_i \times \mathbf{V}_i) \cdot \mathbf{Y} \right] \right\} \quad (20a)$$

when $Y^2 < \infty$, and

$$\phi(\mathbf{R}) = \sum_{i=1}^n \left\{ V_i^2 + W_i^2 - 2 \left[\mathbf{V}_i \cdot \mathbf{W}_i + 2(\mathbf{V}_i \cdot \mathbf{X})(\mathbf{W}_i \cdot \mathbf{X}) \right] \right\} \quad (20b)$$

where $\mathbf{X}^2 = 1$, when $Y^2 = \infty$.

In the first case, we have

$$\frac{\partial \phi}{\partial y_j} = \frac{4}{(1+Y^2)^2} \sum_{i=1}^n \left\{ 2 \left[(\mathbf{V}_i \times \mathbf{W}_i) \cdot \mathbf{Y} - (\mathbf{V}_i \cdot \mathbf{Y})(\mathbf{W}_i \cdot \mathbf{Y}) - \mathbf{V}_i \cdot \mathbf{W}_i \right] y_j + (1+Y^2) \left[(\mathbf{V}_i \cdot \mathbf{Y}) w_{ij} + (\mathbf{W}_i \cdot \mathbf{Y}) v_{ij} - (\mathbf{V}_i \times \mathbf{W}_i)_j \right] \right\},$$

where the subscript j refers to the j^{th} component of the vector. Thus a necessary condition for ϕ to have a minimum (at least among the rotations of less than 180°) at some Y is given by the vector equation

$$2 \left[\sum (\mathbf{W}_i \times \mathbf{V}_i) \cdot \mathbf{Y} + (\mathbf{V}_i \cdot \mathbf{Y})(\mathbf{W}_i \cdot \mathbf{Y}) + \mathbf{V}_i \cdot \mathbf{W}_i \right] Y = (1+Y^2) \sum \left[(\mathbf{V}_i \cdot \mathbf{Y}) \mathbf{W}_i + (\mathbf{W}_i \cdot \mathbf{Y}) \mathbf{V}_i + \mathbf{W}_i \times \mathbf{V}_i \right] \quad (21)$$

(For convenience, the range of summation is omitted hereinafter; always the letter i will be used for the summation index.)

Taking the dot product of each side with \mathbf{Y} , we obtain

$$2 \sum (\mathbf{V}_i \cdot \mathbf{Y})(\mathbf{W}_i \cdot \mathbf{Y}) = \sum \left[(\mathbf{W}_i \times \mathbf{V}_i) \cdot \mathbf{Y} + 2\mathbf{V}_i \cdot \mathbf{W}_i \right] Y^2 - \sum (\mathbf{W}_i \times \mathbf{V}_i) \cdot \mathbf{Y}.$$

When this last expression is substituted in the vector equation and the factor $1 + Y^2$ is divided out, we obtain

$$\sum \left\{ \left[(\mathbf{W}_i \times \mathbf{V}_i) \cdot \mathbf{Y} + 2\mathbf{V}_i \cdot \mathbf{W}_i \right] \mathbf{Y} - \left[(\mathbf{V}_i \cdot \mathbf{Y}) \mathbf{W}_i + (\mathbf{W}_i \cdot \mathbf{Y}) \mathbf{V}_i + \mathbf{W}_i \times \mathbf{V}_i \right] \right\} = 0 \quad (22)$$

The same substitution into equation (20a) gives

$$\phi(\mathbf{R}) = \sum (\mathbf{V}_i^2 + \mathbf{W}_i^2 - 2\mathbf{V}_i \cdot \mathbf{W}_i) - 2\mathbf{Y} \cdot \sum \mathbf{W}_i \times \mathbf{V}_i \quad (23)$$

when \mathbf{Y} satisfies equation (22). Hence, to minimize ϕ we must take the solution of equation (22) which makes $\mathbf{Y} \cdot \sum \mathbf{W}_i \times \mathbf{V}_i$ a maximum.

Equation (22) may be written in matrix notation as follows:

$$(\mathbf{A}^T \mathbf{Y} \mathbf{I} + \mathbf{B}) \mathbf{Y} = \mathbf{A} \quad (24)$$

where \mathbf{I} is the identity matrix, \mathbf{A} is the vector $\sum \mathbf{W}_i \times \mathbf{V}_i$ and \mathbf{B} is a symmetric matrix, with elements

$$b_{jk} = - \sum (v_{ij} w_{ik} + v_{ik} w_{ij}) \quad j \neq k,$$

$$b_{jj} = 2 \sum (v_i \cdot w_i - v_{ij} w_{ij}).$$

(To see this most easily, write out equations (22) and (24) in component form.)

If $\mathbf{A} = 0$, then $\mathbf{Y} = 0$ is the desired solution (there may be other solutions if $\det \mathbf{B} = 0$, but the value of ϕ will be the same for all solutions). If $\mathbf{A} \neq 0$, then the solutions may be obtained as outlined below.

Multiplying each side of the matrix equation by the adjoint of $A^T YI + B$, we obtain

$$\left[\det (A^T YI + B) \right] Y = \left[\text{adj} (A^T YI + B) \right] A ,$$

and a multiplication of this equation by A^T yields the scalar equation

$$\left[\det (A^T YI + B) \right] A^T Y = A^T \left[\text{adj} (A^T YI + B) \right] A .$$

Denoting the scalar $A^T Y$ by λ , the scalar functions $\det (\lambda I + B)$ and $A^T \left[\text{adj} (\lambda I + B) \right] A$ as $f(\lambda)$ and $g(\lambda)$, respectively, the above scalar equation may be written as

$$\lambda f(\lambda) - g(\lambda) = 0 .$$

Note that the left-hand side, say $h(\lambda)$, is a fourth-degree polynomial in λ , and that $-f(-\lambda)$ is just the characteristic polynomial of the symmetric matrix B .

The solutions to equation (24) are obtained by determining the zeros of $h(\lambda)$ and solving the resulting linear equations. However, in the discussion following equation (23), it was determined that the maximum value of $Y \cdot \Sigma(W_i \times V_i)$, i.e., the largest zero of $h(\lambda)$ (denote this zero by λ_0), leads to the minimum value of $\phi(Y)$.

Since $h(\lambda)$ is a fourth-degree polynomial, λ_0 may be obtained analytically; however, a numerical iterative solution is probably more practical, since the zeros of $h(\lambda)$ are easily bounded. We note, first of all, that $\phi(Y)$ is a non-negative function; therefore, equation (23) implies

$$\lambda_0 \leq \frac{1}{2} \sum_{i=1}^n (V_i - W_i)^2 ,$$

which provides an upper bound.

Since B is symmetric, there exists an orthogonal matrix, P say, such that $P^{-1} B P$ is diagonal. Let Y' and U be vectors such that $Y = P Y'$ and $A = P U$. Then

in terms of Y' and U , equation (24) becomes

$$(U^T Y' I + B) P Y' = P U ,$$

and a premultiplication by P^{-1} gives

$$(U^T Y' I + D) Y' = U ,$$

where D is a diagonal matrix whose entries are merely the eigenvalues of B . (Without loss of generality, we may assume the eigenvalues to be arranged in increasing order, say $\lambda_1 \leq \lambda_2 \leq \lambda_3$.) Multiplying each side of the last equation by $\text{adj}(U^T Y' I + D)$ and then by U^T yields the scalar equation

$$\det(U^T Y' I + D) U^T Y' = U^T \text{adj}(U^T Y' I + D) U .$$

However, $U^T Y' = A^T Y = \lambda$ and D is diagonal, so that this last equation may be written in the form

$$\lambda(\lambda + \lambda_1)(\lambda + \lambda_2)(\lambda + \lambda_3) = u_1^2(\lambda + \lambda_2)(\lambda + \lambda_3) + u_2^2(\lambda + \lambda_1)(\lambda + \lambda_3) + u_3^2(\lambda + \lambda_1)(\lambda + \lambda_2) .$$

The left-hand side of the above equation is $\lambda f(\lambda)$ and the right-hand side is $g(\lambda)$. We may now easily determine the sign of $g(\lambda)$ when $\lambda = -\lambda_3, -\lambda_2, -\lambda_1$ (assuming $\lambda_1 \leq \lambda_2 \leq \lambda_3$), and thus also the sign of $h(\lambda) = \lambda f(\lambda) - g(\lambda)$ at these points. From these considerations, it can be shown that λ_0 is at least as great as $-\lambda_1$, and also that the largest zero of $g(\lambda)$ (provided $g(\lambda)$ is not identically zero, i.e., $A = 0$) is less than or equal to $-\lambda_1$. Hence, a lower bound of our desired zero (λ_0) is $-\lambda_1$, which can be determined only by solving a cubic. But the largest zero of $g(\lambda)$ is also a lower bound, and this can be found by solving a quadratic.

Thus far, we have obtained the minimum of $\phi(R)$ for rotations of less than 180 degrees. To obtain the minimum among all 180 degree rotations, we use the method of Lagrange for solving extremum problems with a constraint. The necessary conditions become

$$-4 \sum (V_i \cdot X) w_{ij} + (W_i \cdot X) v_{ij} + 2\mu x_j = 0 \quad j = 1, 2, 3 \quad (25)$$

where μ is Lagrange's multiplier and X must be such that $X^2 = 1$. In terms of our previous notation, the above conditions may be collected into the single vector equation

$$\left[\left(\frac{\mu}{2} - 2 \sum v_i \cdot w_i \right) I + B \right] X = 0 ,$$

where B is the symmetric matrix introduced in equation (24). This last equation can be satisfied if and only if μ is a root of the cubic equation

$$\det \left[\left(\frac{\mu}{2} - 2 \sum v_i \cdot w_i \right) I + B \right] = 0 ,$$

but the roots of this equation are given by

$$\frac{\mu_k}{2} - 2 \sum v_i \cdot w_i = -\lambda_k \quad (k = 1, 2, 3)$$

where the λ_k are eigenvalues of B . The condition equations thus become

$$(B - \lambda_k I) X = 0 , \quad X^2 = 1 ,$$

and the solutions are just the unit eigenvectors of B . The value of $\phi(R)$ at each of these solutions may be obtained by multiplying the j^{th} equation of equations (25) by x_j and adding the three equations together. This yields

$$\mu = 4 \sum (v_i \cdot X) (w_i \cdot X) ,$$

which when substituted into the expression for $\phi(\mathbf{R})$ (equation 20b) gives

$$\begin{aligned}\phi(\mathbf{R}) &= \sum (v_i^2 + w_i^2 + 2v_i \cdot w_i) - \mu_k , \\ &= \sum (v_i - w_i)^2 + 2\lambda_k .\end{aligned}$$

Thus the eigenvector (with unit length) corresponding to the minimum eigenvalue (λ_1) gives the minimum of $\phi(\mathbf{R})$ for all rotations of 180 degrees.

The minimum of $\phi(\mathbf{R})$ for rotations other than 180 degrees is given by equation (23), and is

$$\phi(\mathbf{R}) = \sum (v_i - w_i)^2 - 2\lambda_0 ,$$

and it was also shown that $\lambda_0 \geq -\lambda_1$. Hence, the rotation giving the minimum $\phi(\mathbf{R})$ among all rotations is given by

$$\mathbf{Y} = (\lambda_0 \mathbf{I} + \mathbf{B})^{-1} \mathbf{A}, \text{ when } f(\lambda_0) \neq 0 ,$$

or by a 180-degree rotation with axis of rotation \mathbf{X} defined by

$$(\lambda_0 \mathbf{I} + \mathbf{B}) \mathbf{X} = 0 \text{ and } \mathbf{X}^2 = 1 ,$$

when $f(\lambda_0) = 0$, where λ_0 is the largest zero of $h(\lambda)$.

To avoid inverting a near-singular matrix and dealing with large values of the components of \mathbf{Y} when $f(\lambda_0)$ is near zero, the \mathbf{Z} vector representation of \mathbf{R} may

be obtained from equation (17) and the relationship

$$(\lambda_0 \mathbf{I} + \mathbf{B})^{-1} = \frac{\text{adj}(\lambda_0 \mathbf{I} + \mathbf{B})}{\det(\lambda_0 \mathbf{I} + \mathbf{B})} = \frac{\text{adj}(\lambda_0 \mathbf{I} + \mathbf{B})}{f(\lambda_0)}.$$

This gives

$$\mathbf{Z} = \frac{\text{sgn}[f(\lambda_0)]}{\sqrt{[f(\lambda_0)]^2 + \mathbf{A}^T [\text{adj}(\lambda_0 \mathbf{I} + \mathbf{B})]^2 \mathbf{A}}} [\text{adj}(\lambda_0 \mathbf{I} + \mathbf{B})] \mathbf{A}, \text{ when } f(\lambda_0) \neq 0,$$

and

$$(\lambda_0 \mathbf{I} + \mathbf{B}) \mathbf{Z} = \mathbf{0} \text{ and } \mathbf{Z}^2 = \mathbf{1}, \text{ when } f(\lambda_0) = 0,$$

where

$$\text{adj}(\lambda \mathbf{I} + \mathbf{B}) = \lambda^2 \mathbf{I} + \lambda [\text{Tr}(\mathbf{B}) \mathbf{I} - \mathbf{B}] + \text{adj}(\mathbf{B}),$$

$$f(\lambda) = \lambda^3 + \text{Tr}(\mathbf{B}) \lambda^2 + \text{Tr}(\text{adj}(\mathbf{B})) \lambda + \det(\mathbf{B}),$$

$$g(\lambda) = \mathbf{A}^T [\text{adj}(\lambda \mathbf{I} + \mathbf{B})] \mathbf{A},$$

and $\text{Tr}(\mathbf{B})$ denotes the trace of \mathbf{B} . These last equations are easily verified by direct calculations.

If $n = 2$ and $\mathbf{V}_1^2 = \mathbf{V}_2^2 = \mathbf{W}_1^2 = \mathbf{W}_2^2$, we conjecture that the following simple procedure gives the least squares solution. Let

$$\mathbf{U}_1 = \frac{\mathbf{V}_1 \times \mathbf{V}_2}{|\mathbf{V}_1 \times \mathbf{V}_2|}, \quad \mathbf{U}_1' = \frac{\mathbf{W}_1 \times \mathbf{W}_2}{|\mathbf{W}_1 \times \mathbf{W}_2|},$$

$$U_2 = \frac{V_1 + V_2}{|V_1 + V_2|}, \quad U_2' = \frac{W_1 + W_2}{|W_1 + W_2|},$$

and obtain the rotation which takes each U_i into U_i' , by the techniques of the preceding section. Such a rotation takes the plane determined by V_1 and V_2 into the plane determined by W_1 and W_2 and insures that $W_1 \cdot RV_1 = W_2 \cdot RV_2$. That such a rotation is the least squares rotation has not been verified by direct calculation due to the lengthy and laborious algebra involved; however, symmetry considerations allow no different solution to be proposed.

We have also been very successful in solving equation (21) in the cases $n > 2$ by successive substitutions, i.e., using the iteration

$$Y_{j+1} = \frac{1 + Y_j^2}{2 \sum_{i=1}^n [(W_i \times V_i) \cdot Y_j + (V_i \cdot Y_j)(W_i \cdot Y_j) + V_i \cdot W_i]} \sum_{i=1}^n [(V_i \cdot Y_j) W_i + (W_i \cdot Y_j) V_i + W_i \times V_i].$$

Here, Y_j is the j^{th} approximation to Y , and Y_0 is obtained as above using two of the V_i and their corresponding images W_i . In fact, in one case studied, the procedure converged even when the angle of rotation was as large as 179 degrees.

7. RELATIONS BETWEEN EULERIAN ANGLES, MATRIX OF ROTATION AND VECTOR APPROACH

Since Eulerian angles have a wide usage (especially when the angles correspond to yaw, pitch, and roll) it may be convenient or necessary to transform the matrix or vector parametrization of a rotation into Eulerian angles. To do this, a convention or positive sense of rotation must be established. Here, we assume that the matrices of the simple rotations about each of the coordinate axes are given by

$$R_1(\theta) = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta & \sin \theta \\ 0 & -\sin \theta & \cos \theta \end{pmatrix}, \quad R_2(\theta) = \begin{pmatrix} \cos \theta & 0 & -\sin \theta \\ 0 & 1 & 0 \\ \sin \theta & 0 & \cos \theta \end{pmatrix}$$

$$R_3(\theta) = \begin{pmatrix} \cos \theta & \sin \theta & 0 \\ -\sin \theta & \cos \theta & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

where $R_i(\theta)$ indicates a rotation of θ about the i^{th} axis. The matrix of any rotation R can then be written as

$$R = R_k(\theta_3) R_j(\theta_2) R_i(\theta_1)$$

$$\begin{aligned} i &= 1, 2, 3 \\ j &= 1, 2, 3 \\ k &= 1, 2, 3 \\ i &\neq j \neq k \end{aligned}$$

If i , j , and k are distinct, then by direct calculation one finds that

$$\sin \theta_2 = \delta_{ijk} r_{ki} \quad , \quad \cos \theta_2 \sin \theta_1 = -\delta_{ijk} r_{kj} \quad ,$$

$$\cos \theta_2 \cos \theta_1 = r_{kk} \quad , \quad \cos \theta_2 \sin \theta_3 = -\delta_{ijk} r_{ji} \quad ,$$

$$\cos \theta_2 \cos \theta_3 = r_{ii} \quad ,$$

where $\delta_{ijk} = 1$ if ijk is a cyclic permutation of 123 and $\delta_{ijk} = -1$ otherwise. Thus if R is given, the angles are defined as follows:

$$\sin \theta_2 = \delta_{ijk} r_{ki} \quad ,$$

$$\cos \theta_2 = \pm \sqrt{r_{ii}^2 + r_{ji}^2} = \pm \sqrt{r_{kk}^2 + r_{kj}^2} \quad ;$$

if $\cos \theta_2 \neq 0$,

$$\sin \theta_1 = -\delta_{ijk} r_{kj} / \cos \theta_2 \quad , \quad \cos \theta_1 = r_{kk} / \cos \theta_2 \quad ,$$

$$\sin \theta_3 = -\delta_{ijk} r_{ji} / \cos \theta_2 \quad , \quad \cos \theta_3 = r_{ii} / \cos \theta_2$$

if $\cos \theta_2 = 0$, θ_1 and θ_3 are subject only to the conditions

$$\sin (\theta_3 \pm \delta_{ijk} \theta_1) = \delta_{ijk} r_{ij} ,$$

$$\cos (\theta_3 \pm \delta_{ijk} \theta_1) = \mp \delta_{ijk} r_{ik} ,$$

where the upper signs are taken if $\sin \theta_2 = 1$ and the lower signs when $\sin \theta_2 = -1$.

The factorization is not unique even when $\cos \theta_2 \neq 0$ since either choice of sign for $\cos \theta_2$ produces the same product matrix R.

To factor R in the form

$$R = R_i (\theta_3) R_j (\theta_2) R_i (\theta_1)$$

(where the first and last factors are of the same form), let

$$\sin \theta_2 = \pm \sqrt{r_{ij}^2 + r_{ik}^2} = \pm \sqrt{r_{ji}^2 + r_{ki}^2} ,$$

$$\cos \theta_2 = r_{ii} ;$$

if $\sin \theta_2 \neq 0$,

$$\sin \theta_1 = r_{ij} / \sin \theta_2 , \quad \cos \theta_1 = \delta_{ji} r_{ik} / \sin \theta_2 ,$$

$$\sin \theta_3 = r_{ji} / \sin \theta_2 , \quad \cos \theta_3 = -\delta_{ji} r_{ki} / \sin \theta_2 ,$$

where $\delta_{ji} = 1$ if ji is in natural cyclic order and $\delta_{ji} = -1$ otherwise. If $\sin \theta_2 = 0$, θ_1 and θ_3 are only subject to the conditions

$$\sin (\theta_3 \pm \theta_1) = \delta_{ji} r_{kj} ,$$

$$\cos(\theta_3 \pm \theta_1) = \pm r_{kk} ,$$

where the plus sign is taken if $r_{ii} > 0$ and the minus sign if $r_{ii} < 0$.

Given Eulerian angles $\theta_1, \theta_2, \theta_3$ as defined above, the vector representation can be obtained by forming the product matrix R , and then using the techniques of the previous sections. A more direct approach, however, is to use equations (15) or (16) twice, where

$$Y_i = \tan \frac{\theta}{2} E_i , \quad Z_i = \sin \frac{\theta}{2} E_i , \quad -\pi < \theta \leq \pi ,$$

are the vectors corresponding to $R_i(\theta)$ and E_i is the coordinate axis about which the rotation is taken (if $\pi < \theta < 2\pi$ then the negatives of the above expressions for the range $-\pi < \theta < 0$ must be used).

To obtain the Eulerian angles from the vector Z , we have from equation (5)

$$r_{ij} = 2 \left[z_i z_j \pm z_k \sqrt{1 - Z^2} \right] , \quad i \neq j$$

$$r_{ii} = 1 - 2Z^2 + 2z_i^2 ,$$

where the upper sign is used when ij is in natural cyclic order and the lower sign otherwise. The angles are then obtained from the appropriate formulas above.

8. EQUATIONS OF MOTION OF A RIGID BODY

If $R(t)$ is the matrix of a rotation which defines the orientation of a coordinate system (attached to a moving rigid body) relative to a fixed coordinate system then R satisfies the matrix differential equation

$$\dot{R}(t) = \Omega(t) R(t) , \quad R(0) = I , \quad (26)$$

where Ω is a skew-symmetric matrix such that $\Omega V = V \times \omega(t)$ for all vectors V , and ω is the angular velocity vector. Writing this equation in the form

$$\dot{R}R^{-1} = \Omega ,$$

and letting $R = R_k(\theta_3) R_j(\theta_2) R_i(\theta_1)$, we obtain

$$\Omega = \dot{R}R^{-1} = R_k R_j \dot{R}_i R_i^{-1} R_j^{-1} R_k^{-1} + R_k \dot{R}_j R_j^{-1} R_k^{-1} + \dot{R}_k R_k^{-1} . \quad (27)$$

Although this is a matrix equation, it represents only three independent component equations, since each of the product matrices on the right is skew-symmetric. These three independent equations can be collected into a single vector equation by using the well-known isomorphism between 3×3 skew-symmetric matrices and three-dimensional vectors,

$$S(V) = \begin{pmatrix} 0 & v_3 & -v_2 \\ -v_3 & 0 & v_1 \\ v_2 & -v_1 & 0 \end{pmatrix} \leftrightarrow \begin{pmatrix} v_1 \\ v_2 \\ v_3 \end{pmatrix} = V . \quad (28)$$

It is easy to verify that if R is a rotation matrix, then $RS(V)R^{-1} \rightarrow RV$. Also, from the definitions of R_i , it is straightforward to show that $\dot{R}_i(\theta)R_i^{-1}(\theta) \rightarrow \theta E_i$, where E_i is the coordinate axis of rotation (this is also valid for rotations about any fixed line). From these considerations, and the fact that $\Omega \leftrightarrow \omega$, equation (27) is equivalent to

$$\begin{aligned} \omega &= \dot{\theta}_1 R_k(\theta_3) R_j(\theta_2) E_i + \dot{\theta}_2 R_k(\theta_3) E_j + \dot{\theta}_3 E_k \\ &= A\dot{\theta} , \quad \theta(0) = 0 , \end{aligned} \quad (29)$$

where A is a matrix whose columns are just the vectors $R_k R_j E_i$, $R_k E_j$, and E_k , respectively, and $\dot{\theta} = (\dot{\theta}_1, \dot{\theta}_2, \dot{\theta}_3)^T$.

The matrix differential equation has no singularities, but requires the integration of nine scalar functions. Equation (29), on the other hand, only involves three scalar functions, but the matrix A is singular when $\cos \theta_2 = 0$ for $k \neq i$, or when $\sin \theta_2 = 0$ for $k = i$. Thus, no set of Eulerian angles can be chosen so that $\dot{\theta}$ will be defined for all rotations. In fact, any set of Eulerian angles gives singularities for rotations as small as ninety degrees.

To obtain the equations of motion expressed in terms of the Y vector, we merely differentiate equation (13), and make the proper substitutions using equations (4) and (13), and the matrix differential equation. This gives

$$\dot{Y} = \frac{1}{2} [(\omega \cdot Y)Y + Y \times \omega + \omega], \quad Y(0) = 0. \quad (30)$$

To solve for ω cross-multiply each side by Y, and subtract the resulting equation from the original. Thus,

$$\frac{1}{2} (1 + Y^2) \omega = \dot{Y} - Y \times \dot{Y}.$$

The differential equation (30) has no singularities, but from the definition of Y, we know that solutions involving 180° rotations will diverge to infinity. For many applications, however, involving only moderate displacements of the moving frame, this will not present any difficulties.

The differential equations in terms of the Z rotation vector can be obtained in a manner similar to that used for the Y rotation vector. However, a more direct approach is to use the identity

$$Z = \frac{1}{\sqrt{1 + Y^2}} Y, \quad (31)$$

whence,

$$\dot{Z} = \frac{1}{\sqrt{1 + Y^2}} \dot{Y} - \frac{Y \cdot \dot{Y}}{(1 + Y^2) \sqrt{1 + Y^2}} Y. \quad (32)$$

From equation (30) we have

$$Y \cdot \dot{Y} = \frac{1}{2} (1 + Y^2) \omega \cdot Y . \quad (33)$$

Thus, combining equations (30) thru (33), we obtain

$$\dot{Z} = \frac{1}{2} \left[Z \times \omega + \sqrt{1 - Z^2} \omega \right], \quad Z(0) = 0 . \quad (34)$$

Conversely, let $Z(t)$ be a differentiable vector function such that over some interval, say $t_0 < t < t_1$, $Z^2 < 1$ and Z satisfies equation (34). Then $Z(t)$ defines a rotation matrix $R(t)$ by equation (5) and \dot{R} is given by

$$\dot{R} = -4Z \cdot \dot{Z}I + 2(\dot{Z}Z^T + Z\dot{Z}^T) + 2\sqrt{1 - Z^2} S(\dot{Z}) - 2Z \cdot \dot{Z}(1 - Z^2)^{-1/2} S(Z)$$

where $S(V)$ denotes the skew-symmetric matrix formed from the vector V by equation (28). Taking the dot product of each side of equation (34) gives

$$Z \cdot \dot{Z} = \frac{1}{2} \frac{d(Z^2)}{dt} = \frac{1}{2} \omega \cdot Z \sqrt{1 - Z^2} ,$$

or

$$Z \cdot \dot{Z}(1 - Z^2)^{-1/2} = \frac{1}{2} \omega \cdot Z , \quad (35)$$

since $Z^2 < 1$. When equations (34) and (35) are substituted into the expression for \dot{R} we find

$$\begin{aligned} \dot{R} = \sqrt{1 - Z^2} \left[Z\omega^T + \omega Z^T + S(Z \times \omega + \sqrt{1 - Z^2} \omega) - 2\omega \cdot ZI \right] + Z(Z \times \omega)^T + (Z \times \omega)Z^T \\ - \omega \cdot ZS(Z) \end{aligned}$$

and direct calculation will indeed verify that $\dot{\mathbf{R}} = \Omega\mathbf{R}$ for $t_0 < t < t_1$. In fact, $\dot{\mathbf{R}} = \Omega\mathbf{R}$ even if $Z^2 = 1$ provided Z has a derivative at this point satisfying equation (34) and equation (35) is valid in the limit as Z^2 approaches unity. For example, if ω is a non-zero constant vector, then

$$\mathbf{Z} = \sin\left(\frac{|\omega| t}{2}\right) \frac{\omega}{|\omega|}$$

is a function satisfying equation (34) for $-\pi \leq |\omega| t \leq \pi$ and indeed the matrix \mathbf{R} defined by \mathbf{Z} satisfies equation (26) over the same interval.

Thus, both equations (30) and (34) define the motion of a rigid body over a wider range of allowable orientations than Euler's equations (equation 29) and only require the integration of three scalar equations which do not contain trigonometric functions. Furthermore, the results of the integration (especially \mathbf{Z}) can be used directly, without having to generate the matrix, e.g. the coordinates of a vector relative to the fixed system may be obtained directly from the coordinates relative to the body system and \mathbf{Y} or \mathbf{Z} per Section 3; if \mathbf{Z}_1 defines the orientation of the body at time t_1 relative to the body system at $t = 0$, and \mathbf{Z}_2 defines the orientation when $t = t_2$ relative to the body system at $t = t_1$, then the rotation product $\mathbf{Z} = \mathbf{Z}_2^* \mathbf{Z}_1$ gives the orientation at time t_2 relative to the body system at $t = 0$ for all \mathbf{Z}_1 and \mathbf{Z}_2 . Eulerian angles can also be obtained as described in the previous section.

As with most differential equations, one would have to devote considerable time to equations (30) and (34) in order to describe completely the properties of the solutions. However, we mention only one common property of both equations, which is useful for approximating the solutions for small increments of time. If $\mathbf{Y}(0) = \mathbf{Z}(0) = 0$, then we note that the n^{th} derivative of \mathbf{Y} or \mathbf{Z} at $t = 0$ contains the term $\omega^{(n-1)}(0)$ (the $(n-1)^{\text{th}}$ derivative of ω at $t = 0$), and for the first two derivatives this is the only term. Thus,

$$\mathbf{Y}(h) = \frac{1}{2} \int_0^h \omega dt + \mathbf{U}$$

$$\mathbf{Z}(h) = \frac{1}{2} \int_0^h \omega dt + \mathbf{V} ,$$

where U and V are of the order of h^3 . Hence, to second order, both solutions may be approximated by the integral of the angular velocity.

9. CONCLUSION

The significant advantages of the vector approach to rotations as presented here over other parametrizations is that the vector parameters can be obtained with ease from basic data, and need not be transformed to a new set in order to perform the algebra of rotations (the product of rotations and the product of a vector by a rotation). Not only is the need for evaluating trigonometric functions removed, but the singularities introduced when one attempts to obtain the polar form of a vector or factor a rotation matrix into simple rotations are also removed. To illustrate these last remarks, we cite one final important application of our vector approach to rotations.

In orbit theory, it is customary to obtain the components of the position and velocity vectors by rotating a coordinate system in which the direction cosines of the angular momentum vector are given by $\mathbf{E}_3 = (0, 0, 1)^T$, into a fixed system in which these direction cosines are also known, say $\mathbf{H} = (h_1, h_2, h_3)^T$. This is usually done by assuming \mathbf{H} is of the form $\mathbf{H} = (\sin i \sin \Omega, -\sin i \cos \Omega, \cos i)^T$; thus the rotation is normally given by $\mathbf{R} = \mathbf{R}_3(-\Omega) \mathbf{R}_1(-i)$. Unfortunately, this technique produces a singularity even in the trivial case $i = 0$ (no rotation required). On the other hand, from Section 4 we find almost immediately that

$$\mathbf{z} = \frac{1}{\sqrt{2(1+h_3)}} \mathbf{H} \times \mathbf{E}_3 = \begin{pmatrix} -\sin \frac{i}{2} \cos \Omega \\ -\sin \frac{i}{2} \sin \Omega \\ 0 \end{pmatrix}$$

is the rotation vector taking \mathbf{E}_3 into \mathbf{H} . Thus, the two parameters z_1 and z_2 define the rotation uniquely for all i except $i = \pi$.

ACKNOWLEDGMENT

The author wishes to express his sincere thanks to Mr. Richard desJardins for his painstaking review of the manuscript and for his many helpful suggestions concerning the manner of presentation.

REFERENCES

1. R. A. Frazer, W. J. Duncan, A. R. Collar, Elementary Matrices, The Macmillan Company, New York, 1947.
2. A. Mayer, "Rotations and Their Algebra," SIAM Review, Vol. 2, No. 2, April 1960, pp. 77-122.
3. J. Stuelpnagel, "On The Parametrization of The Three-Dimensional Rotation Group," SIAM Review, Vol. 6, No. 4, October 1964.